



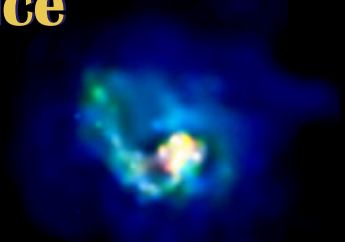
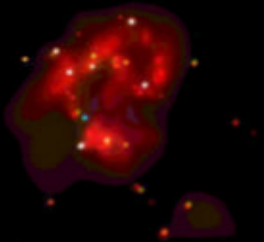
*Con-X FST, November 19-20, 2003*

# Constellation

The Constellation X-ray Mission



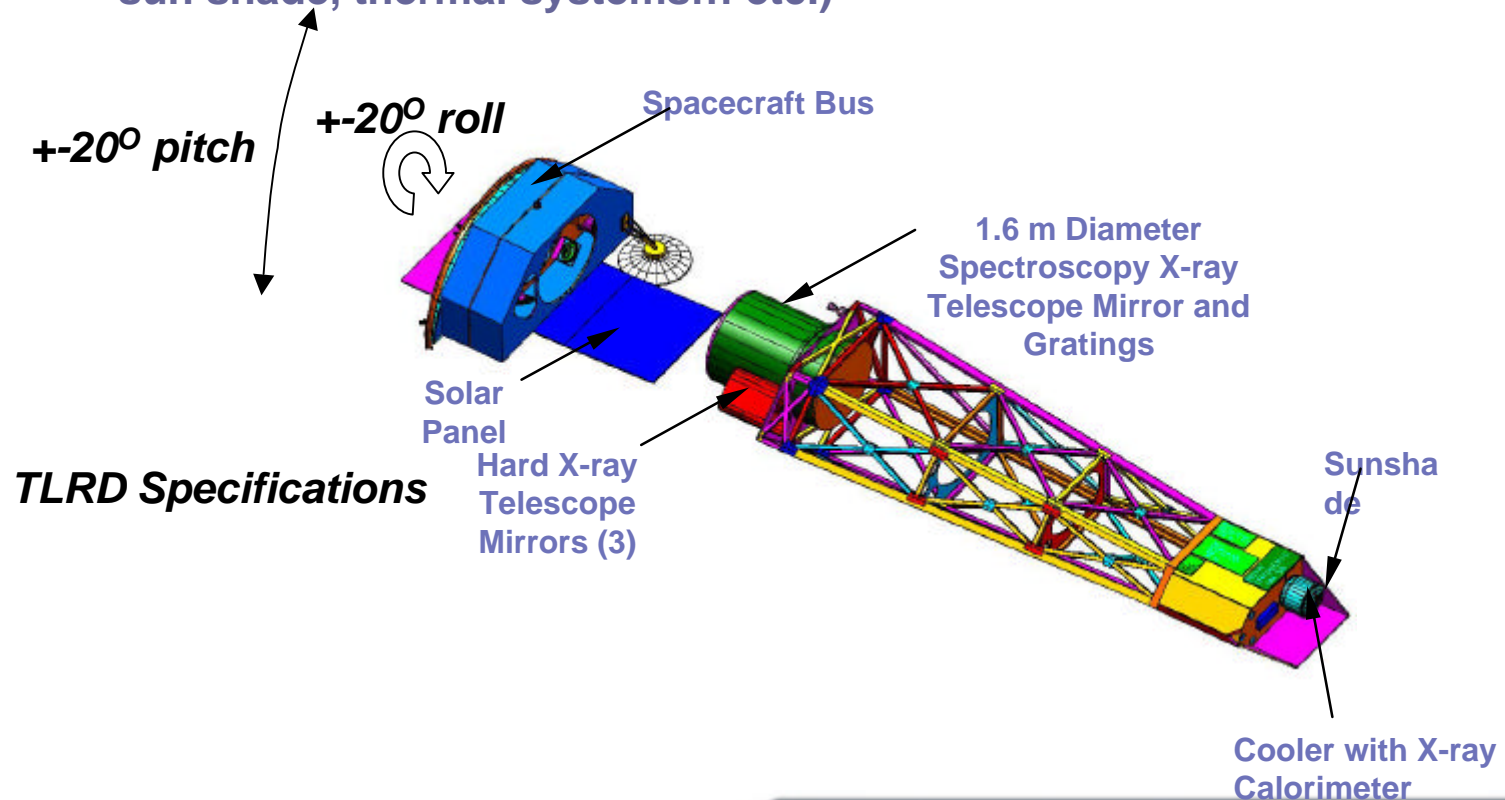
## ►► Sky Coverage Capabilities and ToO Science



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## QUESTION:

- What effects do the current instantaneous sky coverage specs have on ToO driven science?
  - Top Level Requirements Doc,  $\pm 20^\circ$  roll,  $\pm 20^\circ$  pitch = 34% sky (simultaneous)
  - Solar Arrays, Power System sized for this requirement (also, collimators, sun shade, thermal systems... etc.)



## How To Predict Number/Nature of ToOs?

- 1: ODRM – but somewhat incomplete with regards to ToOs in particular
- 2: Chandra Observing Catalog – what has been done that is relevant to Con-X? (Spectral Studies, not locations, jets, etc.)
- 3: XMM Observing Catalog – Compare to Chandra

XMM	Chandra
Pitch = $\pm 28^\circ$ ,	Pitch = $45^\circ \rightarrow 180^\circ$ (articulated SAs)
Sky Coverage = 47%	Sky Coverage = 85% (15% sun block)
Rate of ToOs?	Rate of ToOs?

***ASSUME: ToOs uniformly distributed on sky; sky coverage = capacity to observe ToO***

## Numbers/Types of ToOs?

### ■ 1: ODRM

- Relativistic FeK lines in BHXN in outburst (Miller, Garcia).
  - **Expect ~1 per year**, Chandra (AO2-4) rate 1.3/year
- X-raying the Hot IGM (Mathur, PKS2155-304/Blazar ToO?)
  - Chandra Obs of MKN421 (Nicastro) , rate 1/year, but Con-X can do 10s targets, 5-15s OVIII in 200ks. For IGM abundances high flux levels may be needed... **Expect ~ few per year**
  - ToO Observations of nearby SN1a/c (Hughes, Lewin, Kulkarni...)
  - **Expect ~1 per year** (within 10Mpc, Virgo). Chandra (AO3-4) rate = 4/year (within 30Mpc, GRB like)
  - Distant SN – 15'' PSF causes confusion with galaxy
- That's all for ODRM (currently)

## Number/Types of ToOs?

- **2: Chandra ToO Observations (not yet in ODRM)**
  - Accreting MSPs
    - Discovery rate **1/year**, Chandra rate 1/year
  - Magnetars/SGR
    - Small population, Chandra rate **2/year**
  - Spectra of X-ray Rich/OD GRBs
    - Discovery rate high, but trigger in Con-X era? Chandra rate (spectra) = **3.5/year**.
  - CVs in outburst
    - **2.5/year** with Chandra
  - NS Transients throughout decay – continued accretion (abundances?), heating/cooling of crust?
    - Aql X-1, RB, NS/XRT, targets limited – Chandra rate **1/year**
  - Galactic Nova, Super-Soft Sources, others?

## Numbers/Types of ToOs?

- 3: Compare XMM ( $28^\circ = 47\%$ ) and Chandra (=85%)
  - Chandra (AO3,AO4)
    - Total Number of ToOs + DDT = 75/year
    - Total Number of BHXN ToOs = 29/year, 7.5 distinct objects/year (Jets, Positions; only ~1/year for Broad Fe Lines)
  - XMM
    - Total Number of ToOs + DPS =  $68/3.5 \text{ years} = 19.4/\text{year}$
    - Total Number of BHXN ToOs =  $8/3.5 \text{ years}$ , 5 distinct objects/3.5 years
    - Average rate ~1/4 Chandra
- Lower XMM numbers may be partially due to sky coverage.



## Expected Con-X S Numbers of ToOs in 4-year Nominal Life?

- **BHXN Broad Fe Lines:**
  - $\sim 1.4 = 1/\text{year} * 34\% * 4 \text{ years}$
- **Hot IGM (Missing Baryons)**
  - $\sim 1.4 = 1/\text{year} * 34\% * 4 \text{ years}$  FOR ToOs, abundances, do at lower fluxes?
- **Nearby SN:**
  - $\sim 4.4, = 4/\text{year} * 34\% * 4 \text{ years}$ , confusion within Galaxy may limit z
- **Accreting MSPs:**
  - $\sim 1.4 = 1/\text{year} * 34\% * 4 \text{ years}$
- **Magnetars/SGR**
  - $\sim 2.8 = 2/\text{year} * 34\% * 4 \text{ years}$
- **X-ray Rich/OD GRBs**
  - $\sim 4.8 = 3.5/\text{year} * 34\% * 4 \text{ years}$ , but SWIFT may change our targets, trigger in 10y?
- **CVs in Outburst**
  - $\sim 2.7 = 2/\text{year} * 34\% * 4 \text{ years}$
- **Bursters Throughout Decay**
  - $\sim 1.4, 1/\text{years} * 34\% * 4 \text{ years}$

## ANSWER:

- Q: Does the current instantaneous sky coverage spec limit ToO driven science?
- **A: Yes**, for any ToO with  $< 2$  targets/year. MISSION S
  - S  $\sim 1.4$ : Relativistic FeK lines in BHXN, HOT IGM (at highest fluxes), accreting MSPs, Busters throughout decay.
  - S = 3 to 5: Nearby SN, Magnetars, CVs in outburst, OD/X-ray rich GRBs
- Can we increase these S by 2x or more?



## Pitch Limits, Sky Coverage:

Pitch Range	Fraction of Sky	Fraction Solar Power
<b>±20 degrees – spec</b>	<b>34%</b>	<b>94%</b>
<b>±30 degrees</b>	<b>50%</b>	<b>87%</b>
<b>±45 degrees</b>	<b>71%</b>	<b>71%</b>

**Roll Limits:**  $\pm 20^\circ$  – simultaneous with  $\pm 20^\circ$  pitch. At limit of  $20^\circ$  the solar power is down another 6% ( $= 1 - \cos(20)$ ). IF we couple pitch and roll – allowing only 0 roll at maximum pitch – then power considerations alone will allow and **extended pitch range** (ie,  $\arccos(\cos(20) \cdot \cos(20)) = 28^\circ$ )

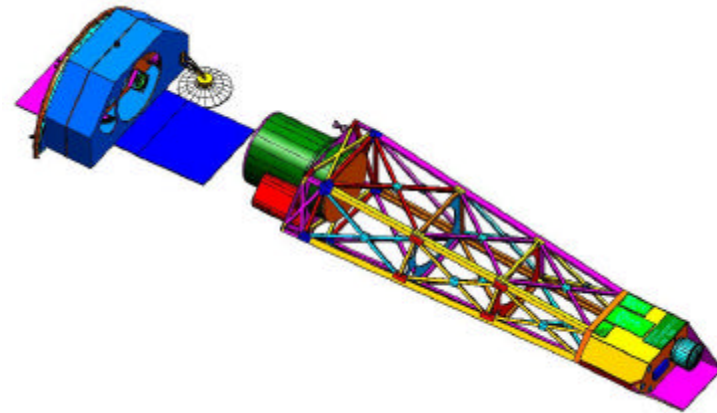
Extended Pitch Range	Fraction of Sky	Fraction Solar Power
<b>±28 degrees – spec?</b>	<b>47%</b>	<b>88%</b>
<b>±35 degrees</b>	<b>56%</b>	<b>77%</b>
<b>±48 degrees</b>	<b>75%</b>	<b>62%</b>

## Impacts/Options of increasing pitch range?

- **Link Pitch and Roll, to allow 28° pitch at 0° roll**
  - ~40% increase in sky coverage, ToO numbers
  - No Impact on SAs, but impacts on Sun Shades, Thermal loading, operations
- **Increase off-nominal Pitch angle to 48° (or intermediate)**
  - 2.2x increase in sky coverage, ToO numbers
  - Require larger SAs, different shape (articulated? \$\$)
    - Modest increase? 88% power → 62% power = 42% SA increase
  - Power System upgrades
  - Current SA size << than SpaceCraft size

## Expected S Numbers of ToOs with $48^\circ$ pitch = 2.2x

- BHXN Broad Fe Lines:
  - ~1.4 -> 3.1
- Hot IGM (Missing Baryons)
  - ~1.4 -> 3.1
- Nearby SN:
  - ~4.4, -> 9.7
- Accreting MSPs:
  - ~1.4 -> 3.1
- Magnetars/SGR
  - ~2.8 -> 6.2
- X-ray Rich/OD GRBs
  - ~4.8 -> 10.6
- CVs in Outburst
  - ~2.7 -> 5.9
- Bursters Throughout Decay
  - ~1.4, -> 3.1



**Significant Increase for these rare ToOs with  
42% increase in SA size**

**End of presentation – backup slides follow**

## Number/Types of ToOs?

- **2: Chandra ToO Observations (not yet in ODRM)**
  - Accreting MSPs (Chakrabarty, Markwardt....)
    - Discovery rate 1/year, Chandra rate 1/year
  - Magnetars/SGR (Kouveliotou, Woods....)
    - Small population, Chandra rate 2/year
  - Spectra of X-ray Rich/OD GRBs (Piro, Paerels, Garmire, Harrison, Ricker...)
    - Discovery rate high, but Chandra rate (spectra) = 3.5/year
  - CVs in outburst (Mauche, Mukai, Long, Beardmore....)
    - 2.5/year with Chandra
  - Bursters throughout decay – continued accretion (abundances?), heating/cooling of crust? (Bildsten, Wijnands, et al)
    - Aql X-1, RB, NS/XRT, targets limited – Chandra rate 1/year
  - Galactic Nova, Super-Soft Sources, others?

# Mission Reference Design

